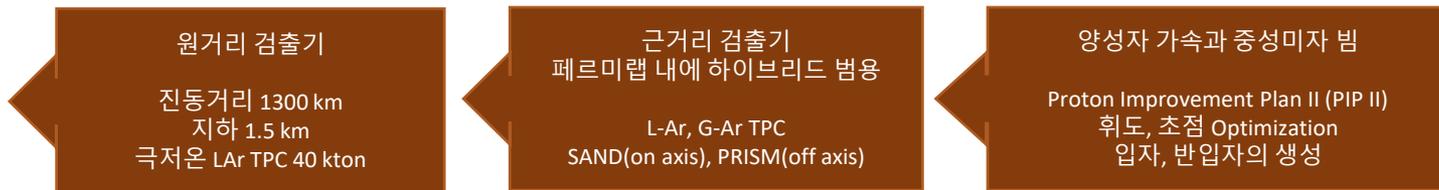


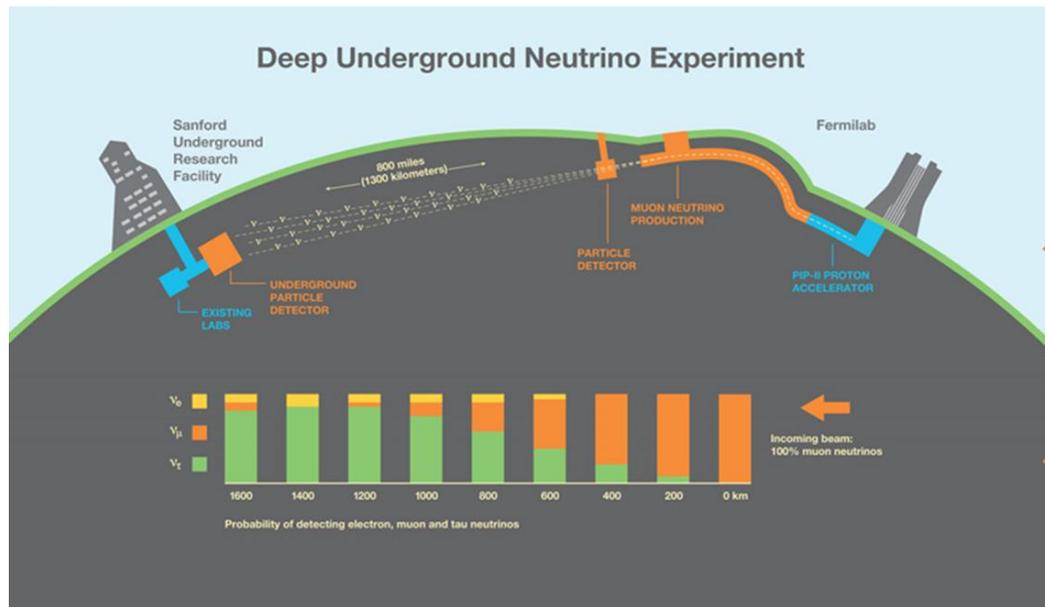
Kim Siyeon  
Chung-Ang University

Nov 18, 2022

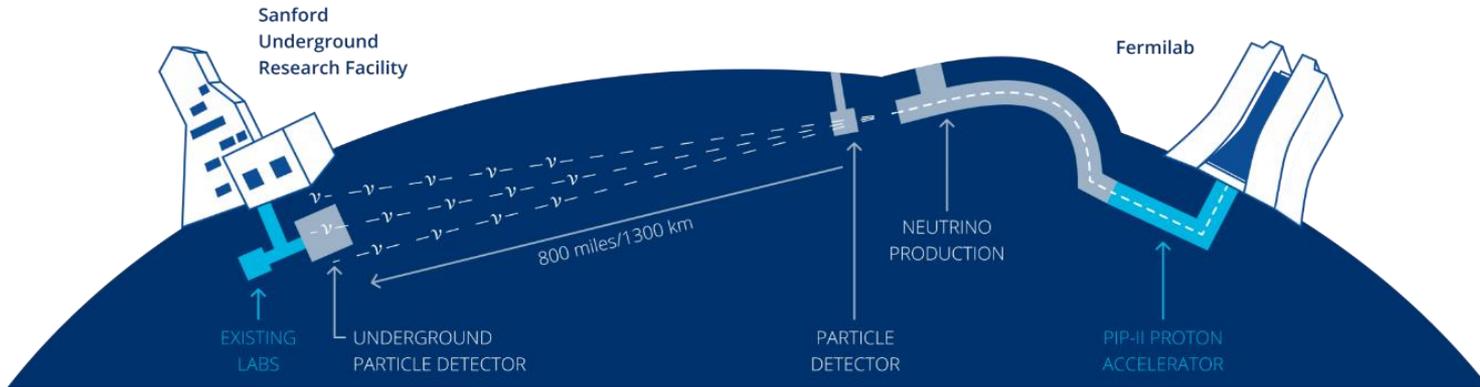
Celebrating the Inauguration of  
Korean Society of High Energy Physics



- 비표준 모형, 암흑물질 탐색
- 우주론과 천체물리  
고에너지 중성미자  
우주 잔재 중성미자  
비활성 중성미자
- 중성미자 질량과 CP  
3세대 중성미자 검증  
CP 비대칭 위상  
질량의 순서
- 중성미자 상호작용  
QE, DIS, RES,  
X-section



- 빔 라인  
진동분석을 위한 개선 (Optimized Beam)
- 근거리 검출기  
진동 전 중성미자 선속 측정,  
오차 개선  
중성미자 상호작용 단면적  
비표준 상호작용 탐색
- 원거리 검출기  
진동효과의 측정  
대기, 태양 중성미자 관측  
초신성 폭발 대기,  
양성자 붕괴



- DUNE Collaboration
- Collaboration Resource Board  
Regina Rameika (Fermilab)
- Institutional Board  
Alfons Weber (Rutherford Lab.)
- DUNE Administration  
Maxine Hronek(Fermilab)
- Fermilab Neutrino Div. Head  
Steve Brice (Fermilab)
- Computing Coordinator  
Heidi Schellman (Fermilab)
- 
- DUNE Cospokesperson  
Stefan Soldner-Rembold (Manchester U.)  
Regina Rameika (Fermilab)

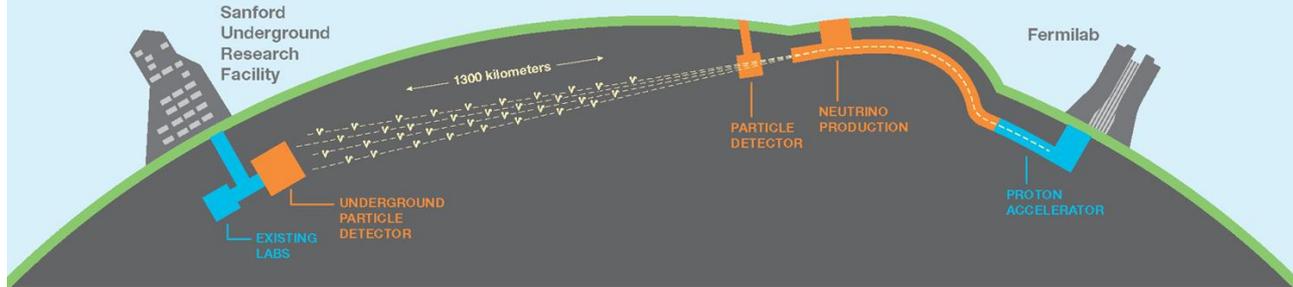
정모세 (UNIST)  
가속기와  
중성미자 빔 인터페이스

김시연(중앙대)  
중성미자 상호작용,  
에너지 재구성, 진동분석

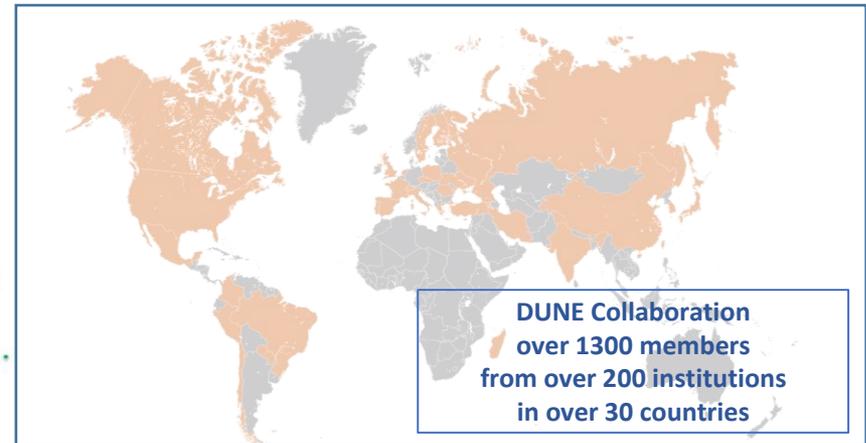
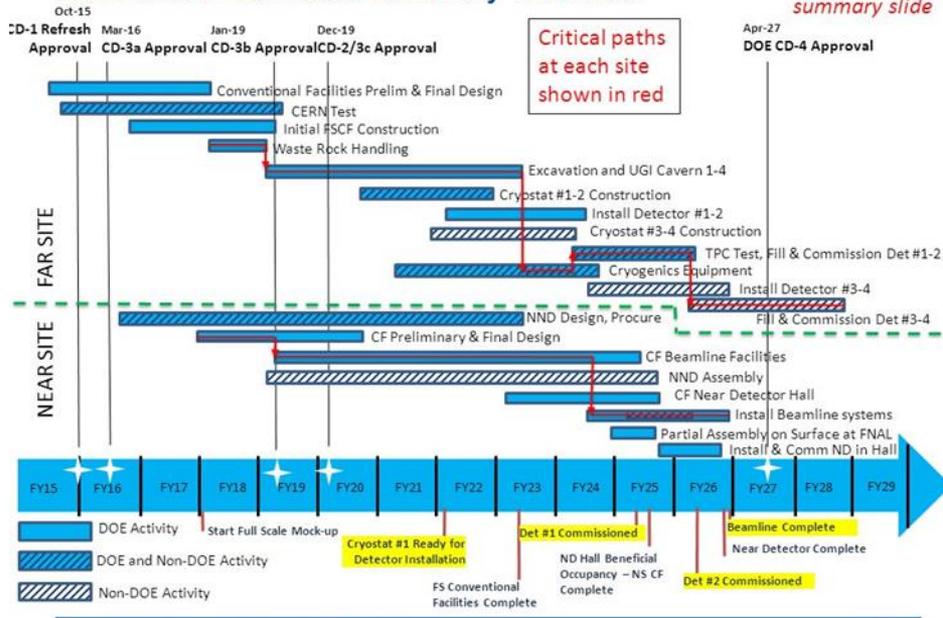
정유선 (중앙대)  
타우중성미자,  
고에너지 중성미자

조기현(KISTI) 신서동 (전북대)  
암흑물질 탐색,  
비표준모형

# Deep Underground Neutrino Experiment

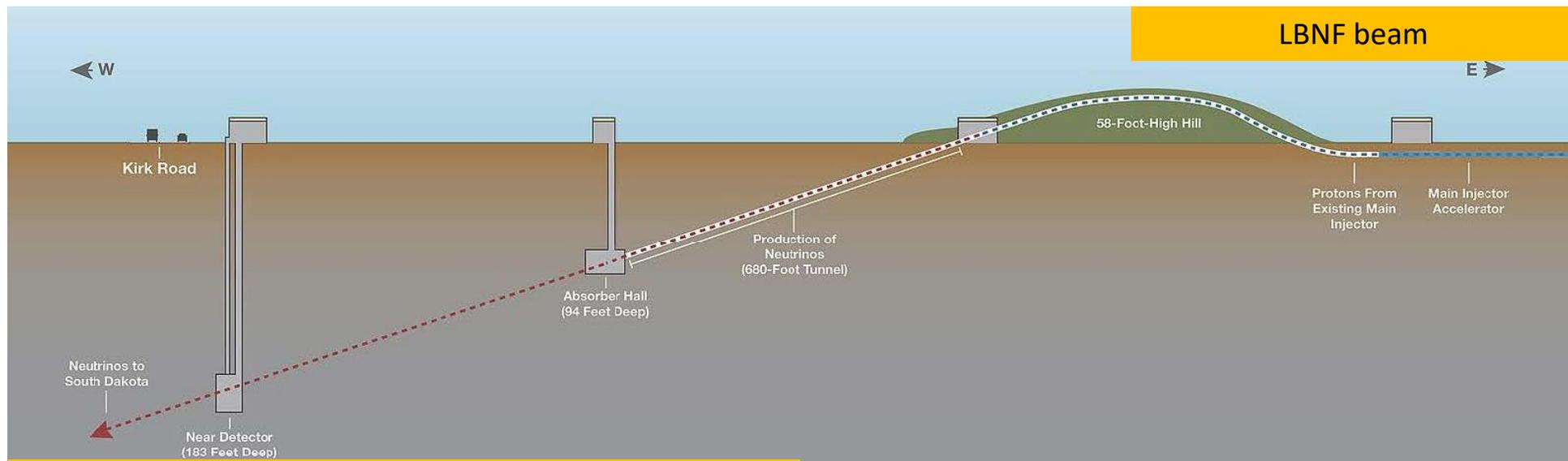


## LBNF/DUNE Schedule Summary Overview



## Korean members of DUNE Collaboration (2022년 10월 현재)

권순우(CAU, Ph.D 학생), 김시연(CAU, IR), 신서동(JBNU, IR), 정유선(CAU, Senior), 정모세(CAU, IR), 조기현(CAU, IR)

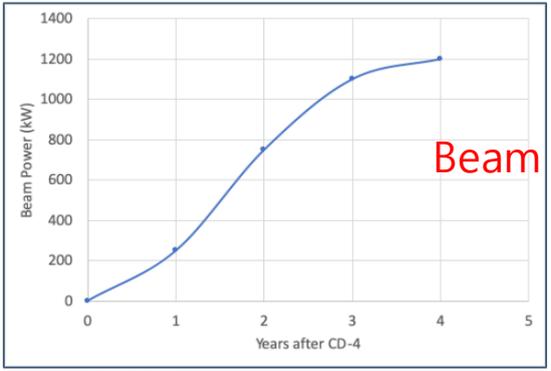
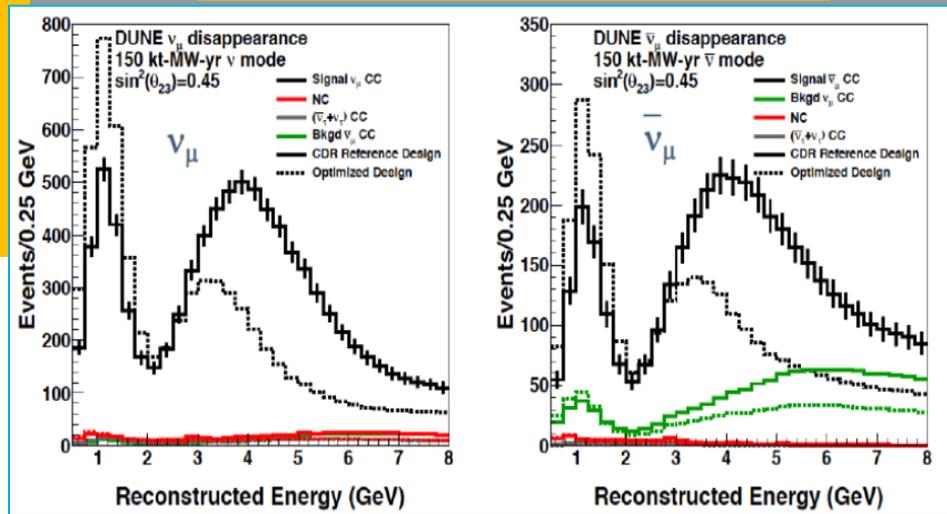


LBNF beam

### Long-Baseline Neutrino Facility

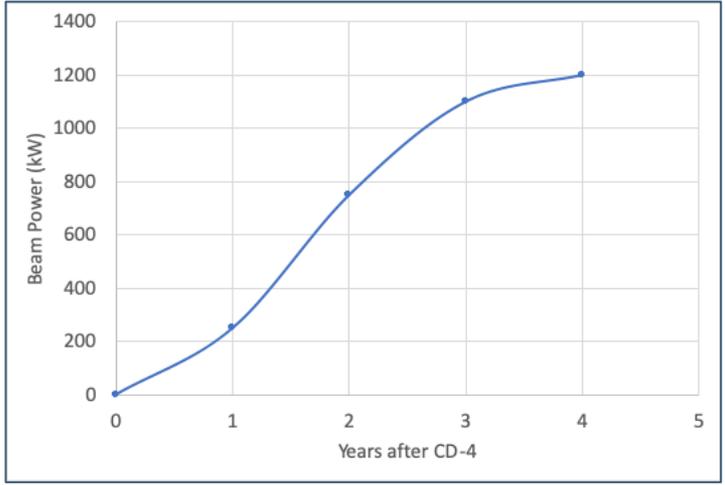
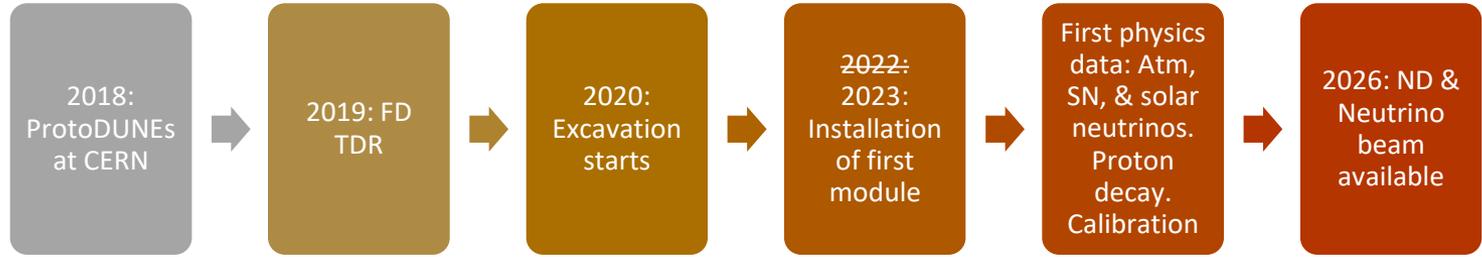
- Proton Improvement Plan (PIP-II)
- Initial 1.2 MW proton beam to be upgraded to 2.4 MW (proton energy 60-120 GeV)
- Beam optimization s. t. more flux at lower energies for better physics sensitivity
- Neutrino beam available in 2026

### Beam Optimization

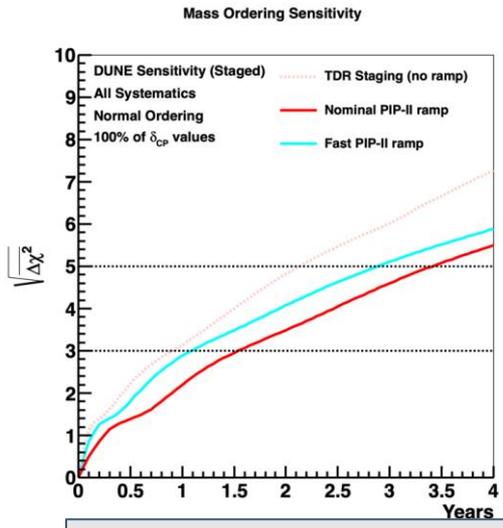


Beam Ramping plan

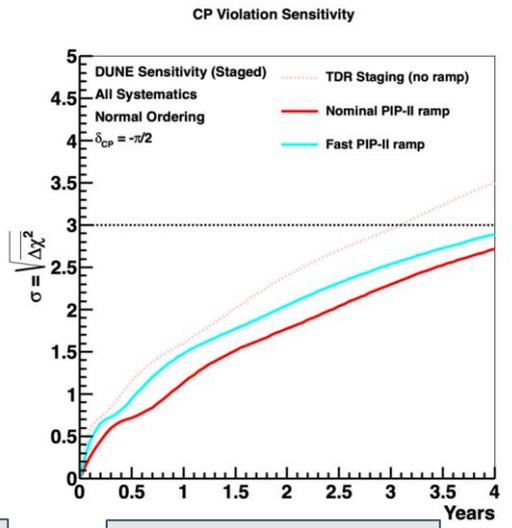
2022-11-18



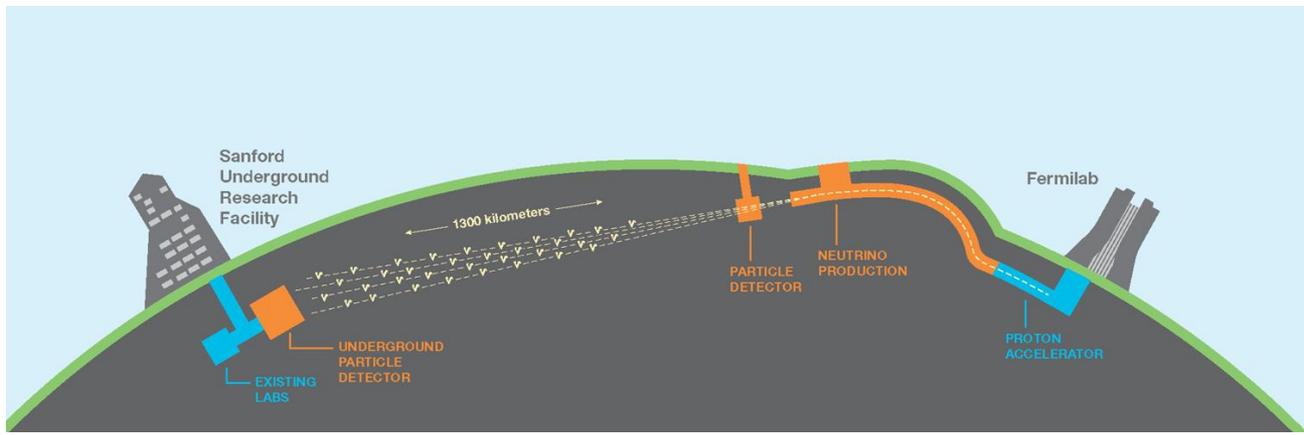
PIP-II Power Ramp



Mass Ordering 100%  $\delta_{CP}$  values



CP Violation  $\delta_{CP} = -\pi/2$

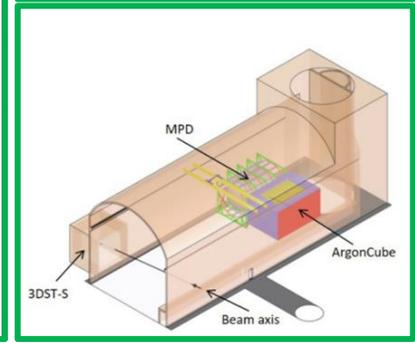
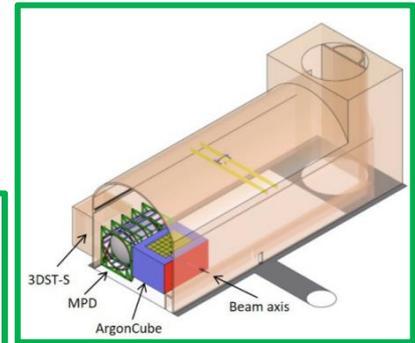
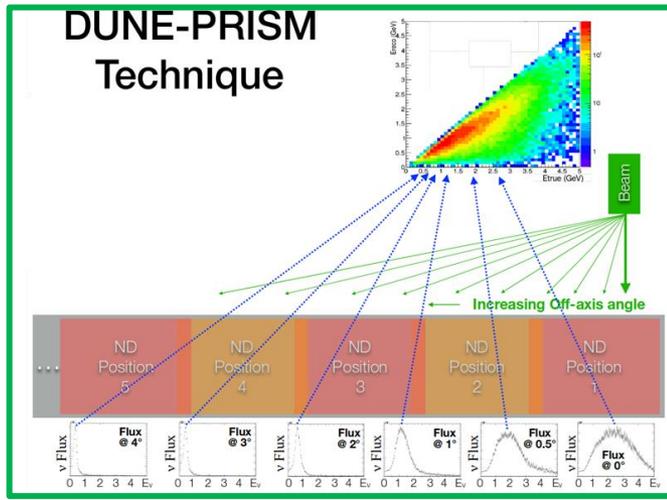
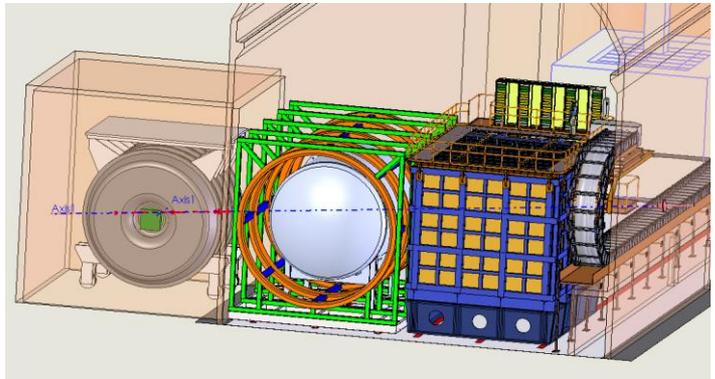


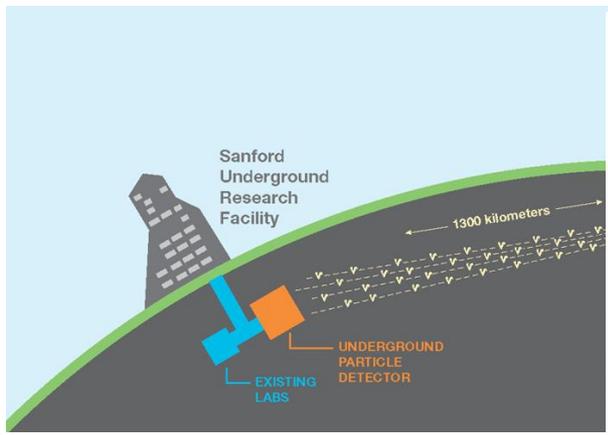
## Near Detector

ND Complex ( 574 m from target hall, 60 m from surface )

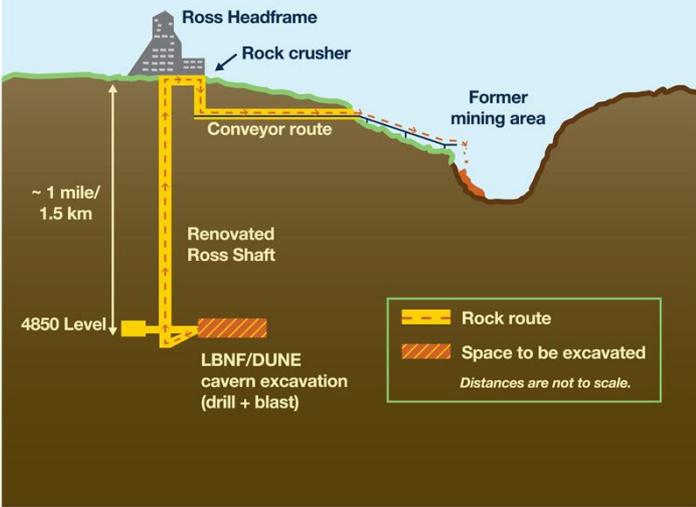
- Liquid Ar Time Projection Chamber
- HP Gas Ar TPC with magnet and ECAL (MPD)
- System for on-Axis Neutrino Detection (SAND)

- Prediction of neutrino flux at FD w/o oscillation
- Control of systematics
- Study neutrino interaction with Ar, CH





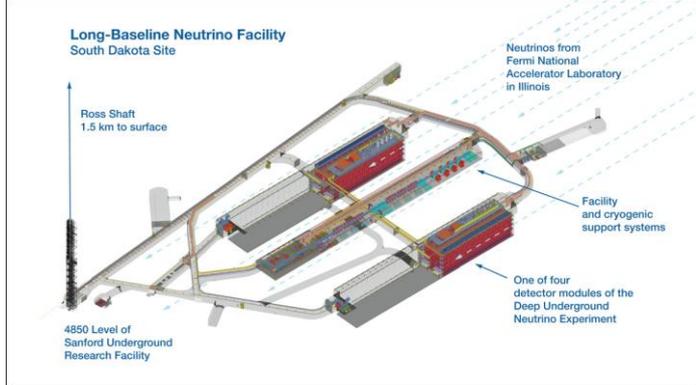
## Excavation of LBNF/DUNE caverns



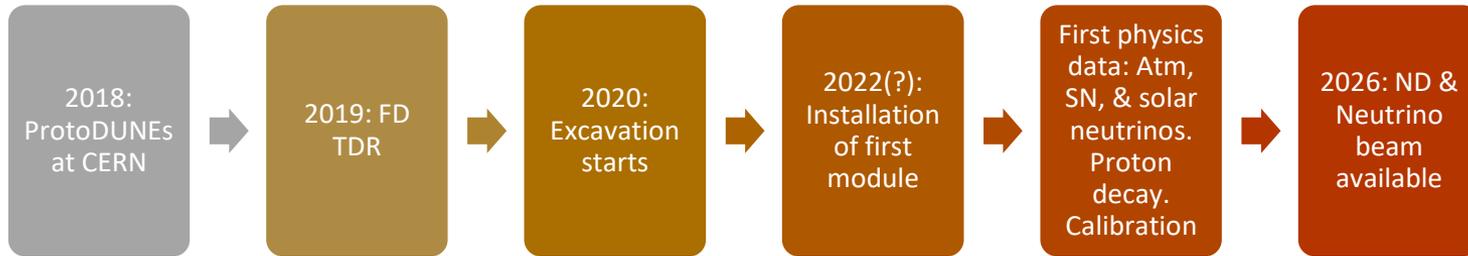
### Far Detector

- Measurement of neutrino events (w/ oscillation)

- #### Non-accelerator Physics Program
- Nucleon decay
  - Atmospheric neutrinos
  - Supernova neutrino bursts
  - Lorentz invariance and CPT violation
  - Astrophysical Neutrinos, e.g., solar neutrinos, diffuse supernove background, and etc.



- 40-kt Liquid Ar time projection chamber (4 x 10 kt)
- 4850 level (4300 mwe)
- The largest cryogenic instrument ever (89K)
- ProtoDUNE at CERN
- Single-phase and ~~double-phase~~ detectors, 2+1+1 (-> Vertical-Drift LAr TPC) -> HD and VD
- The first module will be single-phase. The installation begins in 2022.
- Technical Design Report available.



### DUNE Phase I

- 2 Far Detectors : Horizontal Drift (HD) + Vertical Drift (VD) LAr
- Near Detectors : ND LAr + TMS + SAND + PRISM
- 1.2 MW beam power

### DUNE Phase II

- FD3 + FD4
- ND-Gar replaces TMS.
- 2.4 MW beam power

- Staged year
- 1 (2026) with 20 kt-1.2MW
  - 2 (2027) with 30 kt-1.2 MW
  - 4 (2029) with 40 kt-1.2 MW
  - 7 (2032) with 40 kt-2.4 MW

**DUNE Day 1 : When FD1 is filled and turned on, Science begins.**

# Korean DUNE Activities

- 2016.05 CAU joined DUNE Collaboration
- 2017 ~ 2018 ProtoDUNE L-Ar TPC Single Phase Cold Electronics Module test
- 2018 ~ 2021 3DST Working Group,  
3DST (3-dim Scintillator Tracker) for SAND/ND
  - Joint consortium with T2K SuperFGD Group
  - Prototype LANL Neutron beam test 2019 & 2020
- 2022
- 2023.01 ~ ProtoDUNE HD Data Analysis  
ProtoDUNE VD Cold Electronics
- ProtoDUNE II: Closing TCO in 2022.11, filling LAr in early 2023, OPS for 2023.06 to 2024.07

## 중앙대, 미 에너지부 산하 가속기 연구기관 '페르미랩'과 공동연구센터 설립

발행일 : 2020.04.07

기사만 보기

[출소TV] IBM-코로나19 사태에 따른 상담 업무 환경의 변화 (4/24 생방송)



<중앙대 전경>



**INTERNATIONAL  
COOPERATIVE RESEARCH AND DEVELOPMENT AGREEMENT  
FOR  
BASIC SCIENCE COOPERATION  
(HEREINAFTER "CRADA") NO. FRA-2017-0044  
BY AND AMONG  
FERMI RESEARCH ALLIANCE, LLC  
UNDER ITS U.S. DEPARTMENT OF ENERGY CONTRACT  
NO. DE-AC02-07CH11359  
TO MANAGE AND OPERATE  
FERMI NATIONAL ACCELERATOR LABORATORY  
(HEREINAFTER "LABORATORY")  
AND  
CHUNG-ANG UNIVERSITY**

FOR LABORATORY:

*Nigel S. Lockyer*  
Name: Dr. Nigel S. Lockyer  
Title: Director of Fermilab

Date: November 30, 2018

FOR PARTICIPANT:

*Changsoo Kim*  
Name: Dr. Kim Chang Soo  
Title: President, Chung-Ang University

Date: Nov. 30, 2018





## 최근 중요 실적 및 기여

- 신서동(전북대):  
[Prospects for beyond the Standard Model physics searches at the Deep Underground Neutrino Experiment](#), *Eur.Phys.J.C* 81 (2021) 4, 322,  
 Boosted dark matter search 집필 기여
- 권순우(중앙대):  
[Deep Underground Neutrino Experiment \(DUNE\) Near Detector Conceptual Design Report](#), *Instruments* 5 (2021) 4, 31,  
 Neutron detection from antineutrino events in the 3DST, 분석결과 수록, 집필 기여
- 정기영(중앙대):  
[Muon antineutrino CC 1 neutral pion interaction selection using the invariant mass](#), DUNE-doc-23681-v1, Technical note 작성
- 권순우(중앙대):  
[Neutron detection and application with a novel 3D projection scintillator tracker in the future long-baseline neutrino oscillation experiments](#) Submitted to APC in DUNE

preprint for arXiv

### Neutron detection and application with a novel 3D projection scintillator tracker in the future long-baseline neutrino oscillation experiments

S. Gwon,<sup>1</sup> G. Yang,<sup>2</sup> S. Bolognesi,<sup>3</sup> T. Cai,<sup>4</sup> A. Delbart,<sup>3</sup> A. De Roeck,<sup>5</sup> S. Dolan,<sup>5</sup> G. Eurin,<sup>3</sup> S. Fedotov,<sup>6</sup> G. Fiorentini Aguirre,<sup>7</sup> R. Flight,<sup>4</sup> R. Gran,<sup>8</sup> P. Granger,<sup>3</sup> C. Ha,<sup>1</sup> C.K. Jung,<sup>2</sup> K.Y. Jung,<sup>1</sup> S. Kettell,<sup>9</sup> A. Khotjantsev,<sup>6</sup> M. Kordosky,<sup>10</sup> Y. Kudenko,<sup>6</sup> T. Kutter,<sup>11</sup> J. Maneira,<sup>12</sup> S. Manly,<sup>4</sup> D. Martinez Caicedo,<sup>7</sup> C. Mauger,<sup>13</sup> K. McFarland,<sup>4</sup> C. McGrew,<sup>2</sup> A. Mefodev,<sup>6</sup> O. Mineev,<sup>6</sup> D. Naples,<sup>14</sup> A. Olivier,<sup>4</sup> V. Paolone,<sup>14</sup> S. Prasad,<sup>11</sup> C. Riccio,<sup>2</sup> J. Rodriguez,<sup>7</sup> D. Sgalaberna,<sup>15</sup> A. Sitraka,<sup>7</sup> K. Siyeon,<sup>1</sup> H. Su,<sup>14</sup> A. Teklu,<sup>2</sup> M. Tzanov,<sup>11</sup> E. Valencia,<sup>10</sup> K. Wood,<sup>2</sup> and E. Worcester<sup>9</sup>

## Neutron detection and application with a novel 3D projection scintillator tracker in the future long-baseline neutrino oscillation experiments

- Neutrino-nucleus interaction:  $\nu$ -Ar,  $\nu$ -C,H COH, QE, RES, DIS
- Neutrons in final states: Missing energy in neutrino detection
- Neutron identification:
  - Event-by-event Energy Reconstruction
- 3-dim Scintillator Tracker:
  - DUNE neutrino beam and CH target
  - CCQE-like (cc0pi) event analysis
  - Low- $\nu$  fitting for flux constraint
- LANL neutron beam test (3DST & SuperFGD/T2K):
  - Study of secondary neutrons

S. Gwon,<sup>1</sup> G. Yang,<sup>2</sup> S. Bolognesi,<sup>3</sup> T. Cai,<sup>4</sup> A. Delbart,<sup>3</sup> A. De Roeck,<sup>5</sup> S. Dolan,<sup>5</sup> G. Eurin,<sup>3</sup> S. Fedotov,<sup>6</sup> G. Fiorentini Aguirre,<sup>7</sup> R. Flight,<sup>4</sup> R. Gran,<sup>8</sup> P. Granger,<sup>3</sup> C. Ha,<sup>1</sup> C.K. Jung,<sup>2</sup> K.Y. Jung,<sup>1</sup> S. Kettell,<sup>9</sup> A. Khotjantsev,<sup>6</sup> M. Kordosky,<sup>10</sup> Y. Kudenko,<sup>6</sup> T. Kutter,<sup>11</sup> J. Maneira,<sup>12</sup> S. Manly,<sup>4</sup> D. Martinez Caicedo,<sup>7</sup> C. Mauger,<sup>13</sup> K. McFarland,<sup>4</sup> C. McGrew,<sup>2</sup> A. Mefodev,<sup>6</sup> O. Mineev,<sup>6</sup> D. Naples,<sup>14</sup> A. Olivier,<sup>4</sup> V. Paolone,<sup>14</sup> S. Prasad,<sup>11</sup> C. Riccio,<sup>2</sup> J. Rodriguez,<sup>7</sup> D. Sgalaberna,<sup>15</sup> A. Sitraka,<sup>7</sup> K. Siyeon,<sup>1</sup> H. Su,<sup>14</sup> A. Teklu,<sup>2</sup> M. Tzanov,<sup>11</sup> E. Valencia,<sup>10</sup> K. Wood,<sup>2</sup> and E. Worcester<sup>9</sup>

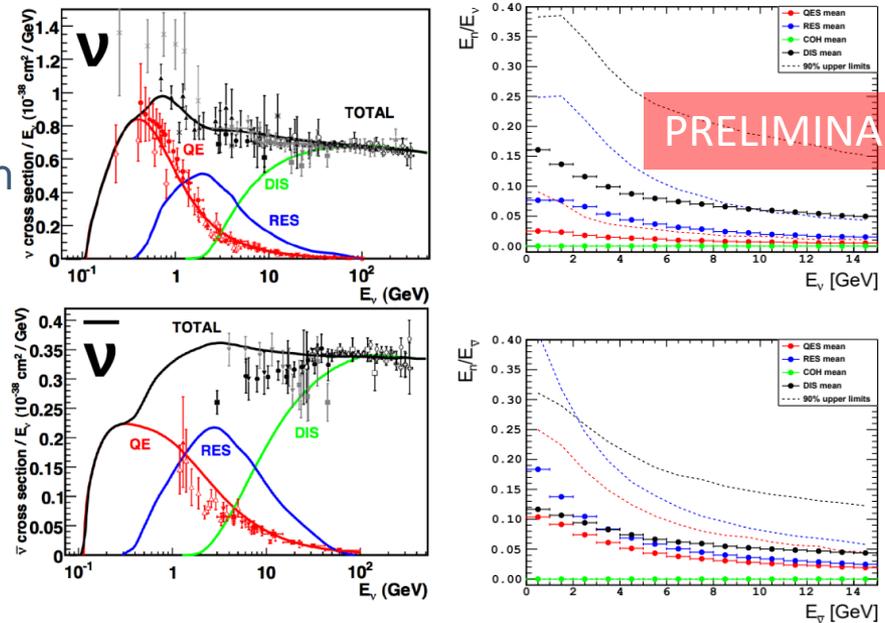
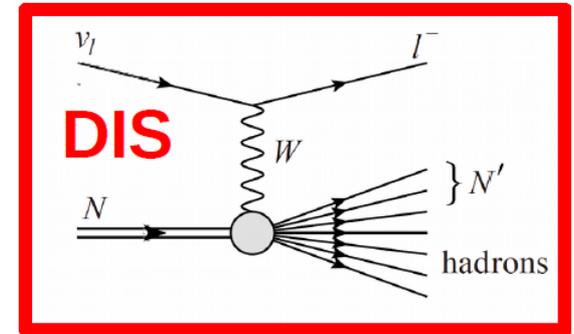
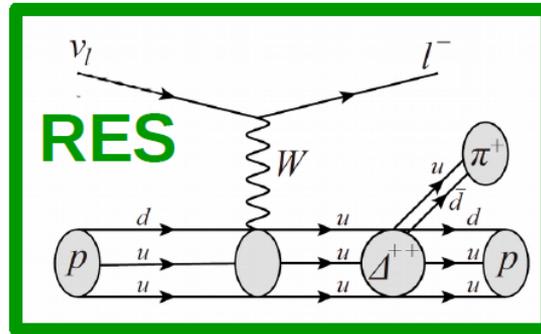
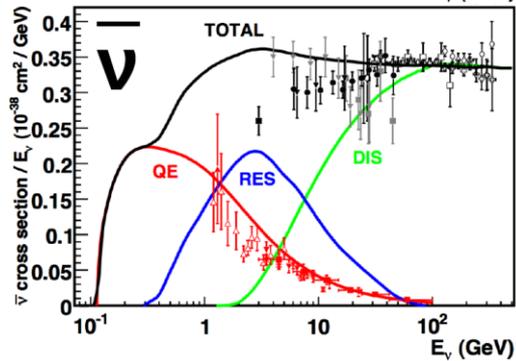
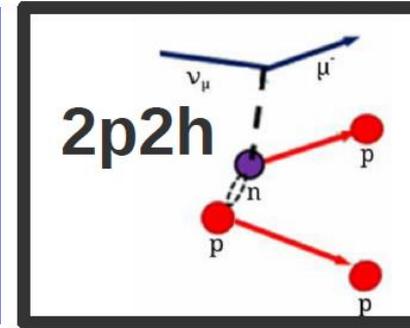
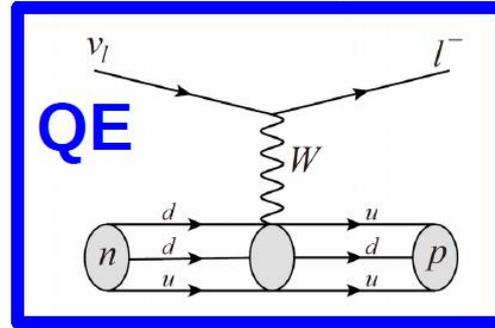
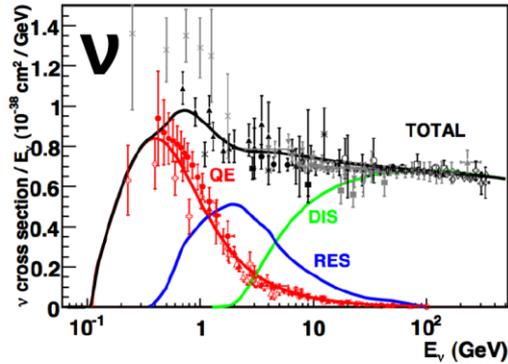


FIG. 1. Average energy fraction delivered to the primary neutrons relative to the neutrino energy (top) and the antineutrino energy (bottom). The average ratios  $E_n/E_\nu$  are in comparison according to the CC Quasi-elastic (QES), CC resonant (RES), CC coherent (COH) and CC deep-inelastic scattering (DIS) interaction modes.

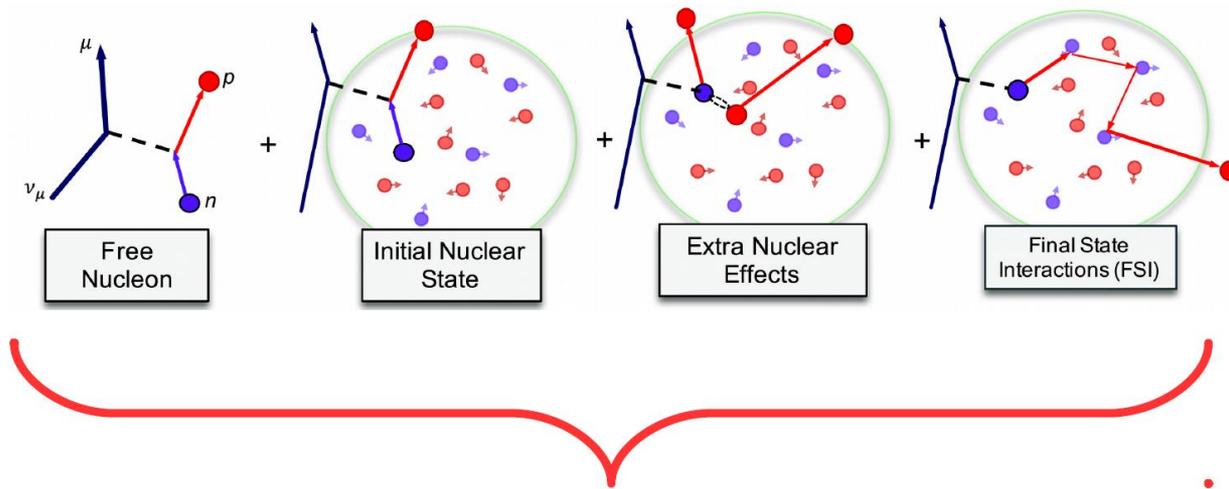
# Neutrino Interaction Physics & the DUNE Near Detector

Mateus F. Carneiro on behalf of the DUNE Collaboration



# Neutrino Interaction Physics & the DUNE Near Detector

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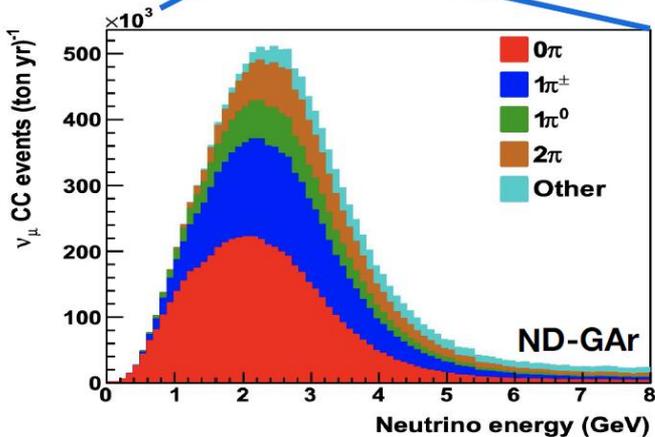
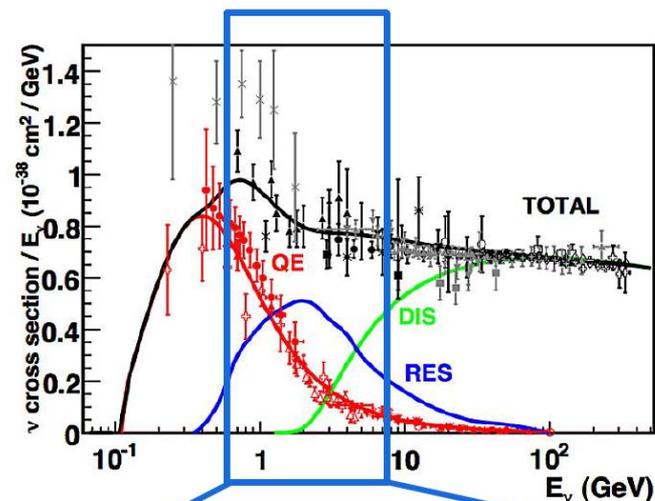


$$R(\vec{x}) = \Phi(E_\nu) \times \sigma(E_\nu, \vec{x}) \times \epsilon(\vec{x}) \times P(\nu_A \rightarrow \nu_B)$$

**Final state particle content does not isolate initial interaction type!**

# Neutrino Interaction Physics & the DUNE Near Detector

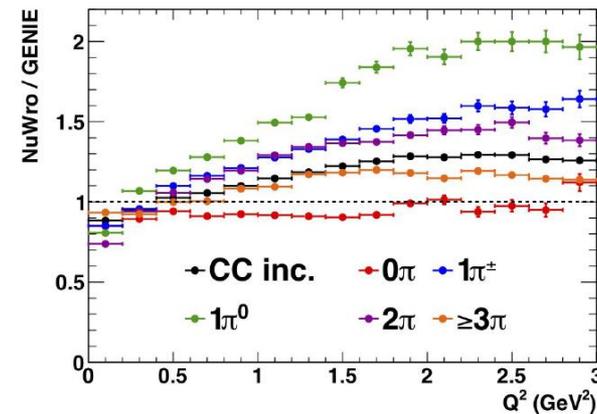
Mateus F. Carneiro on behalf of the DUNE Collaboration



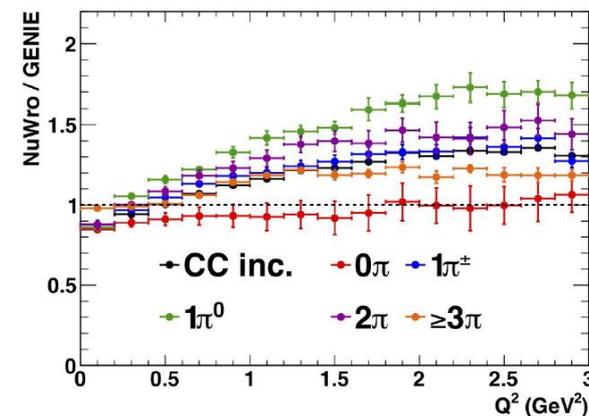
|           |                 | Interaction Channel | Event Rate         |                    |
|-----------|-----------------|---------------------|--------------------|--------------------|
|           |                 |                     | ND-LAr             | ND-GAr             |
| CC        | $\nu_\mu$       |                     | $8.2 \times 10^7$  | $1.64 \times 10^6$ |
|           |                 | $0\pi$              | $2.9 \times 10^7$  | $5.8 \times 10^5$  |
|           |                 | $1\pi^\pm$          | $2.0 \times 10^7$  | $4.1 \times 10^5$  |
|           |                 | $1\pi^0$            | $8.1 \times 10^6$  | $1.6 \times 10^5$  |
|           |                 | $2\pi$              | $1.1 \times 10^7$  | $2.1 \times 10^5$  |
|           |                 | $3\pi$              | $4.6 \times 10^6$  | $9.3 \times 10^4$  |
|           |                 | other               | $9.2 \times 10^6$  | $1.8 \times 10^5$  |
|           | $\bar{\nu}_\mu$ |                     | $3.6 \times 10^6$  | $7.1 \times 10^4$  |
|           | $\nu_e$         |                     | $1.45 \times 10^6$ | $2.8 \times 10^4$  |
| NC        |                 |                     | $5.3 \times 10^5$  | $5.5 \times 10^5$  |
| $\nu + e$ |                 |                     | $8.3 \times 10^3$  | $1.7 \times 10^2$  |

Events per year (1.1×1021 POT)

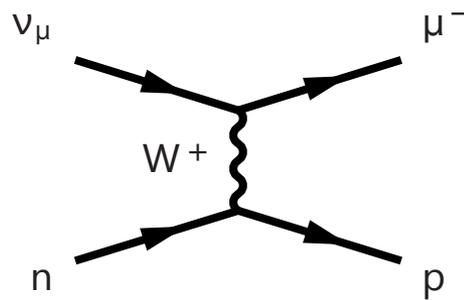
True NuWro/GENIE (FHC)



NuWro/GENIE for various reconstructed final states (FHC)



## Charged-current quasi-elastic scattering - the “golden channel”



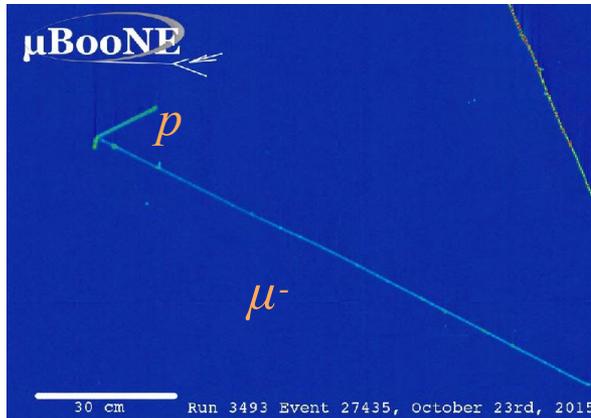
Simple final state - just a muon and a nucleon

Conserve energy and momentum: calculate  $Q^2$  and  $E_\nu$  just from muon kinematics

$$E_\nu = \frac{m_p^2 - (m_n - E_b)^2 - m_\mu^2 + 2(m_n - E_b)E_\mu}{2(m_n - E_b - E_\mu + p_\mu \cos \theta_\mu)}$$

(in a nucleus; binding energy  $E_b = 28$  MeV for argon)

Neutrino mode



Why is this useful?

- Muon has constant  $dE/dx$  (minimum-ionizing particle)
- Long, clear track: **easy to measure  $E_\mu$  and  $\theta_\mu$**
- $\bar{\nu}$  case - neutron hard to detect (neutral)
- Not affected by **final-state interactions**
  - Nucleons can re-interact in the nucleus.

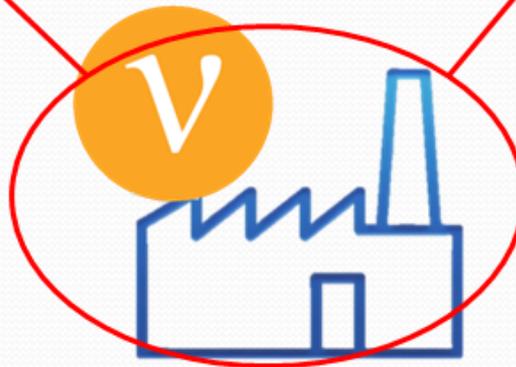
What about anti-muon neutrinos?

What about neutrons?

# BSM Physics in Neutrino Experiments

## Neutrino-sector BSM

- Beyond 3 neutrino flavors (e.g., sterile neutrinos)
- Non-standard interactions of neutrinos
- ...



## Non-neutrino-sector BSM

- (Light) dark matter search
- (Light) mediators or portal scenarios (e.g., dark photon, axion-like particles)
- ...

# DUNE Timeline

## ☐ Phase I

- Two far detectors
- ND-LAr (Movable)
- TMS (Movable) + SAND
- 1.2 MW beam power

Leptophilic mediators

Cosmogenic boosted dark matter (BDM)

## ☐ Phase II

- More far detectors
- ND-GAr (Movable, Replacing TMS)
- 2.4 MW beam power

Axion-like particles (ALP)

## ☐ Proposed off-target mode

Low-mass dark matter (LDM)

# Physics Working Groups

|   |   |  |  |
|---|---|--|--|
| <b>Physics Coordination</b><br>Inés Gil-Botella<br>Chris Marshall | <b>Long-baseline</b><br>Callum Wilkinson<br>Luke Pickering        | <b>High energy</b><br>Lisa Koerner<br>Yun-Tse Tsai   | <b>FD sim/reco</b><br>Chris Backhouse<br>Dom Brailsford      |
| <b>DUNE Physics Working Groups</b>                                | <b>Neutrino Interactions</b><br>Cheryl Patrick<br>Mateus Carneiro | <b>BSM</b><br>Justo Martin-Albo<br>Alex Sousa        | <b>ND sim/reco</b><br>Linda Cremonesi<br>Mat Muether         |
| <b>Liaisons</b><br>Dan Cherdack (ND)<br>Tom Junk (computing)      | <b>Low Energy</b><br>Clara Cuesta<br>Dan Pershey                  | <b>Calibration</b><br>David Caratelli<br>Mike Mooney | <b>protoDUNE analysis</b><br>Leigh Whitehead<br>Tingjun Yang |

Low-energy = 1-10s MeV-scale physics: supernovae, solar, etc.  
This group also works with the backgrounds task force,  
as natural radioactivity is an important background for LE physics

High-energy = GeV-scale non-accelerator physics: atmospheric neutrinos,  
nucleon decay & other signals for which atmospheric neutrinos are a background.  
Formerly known as the nucleon decay WG.

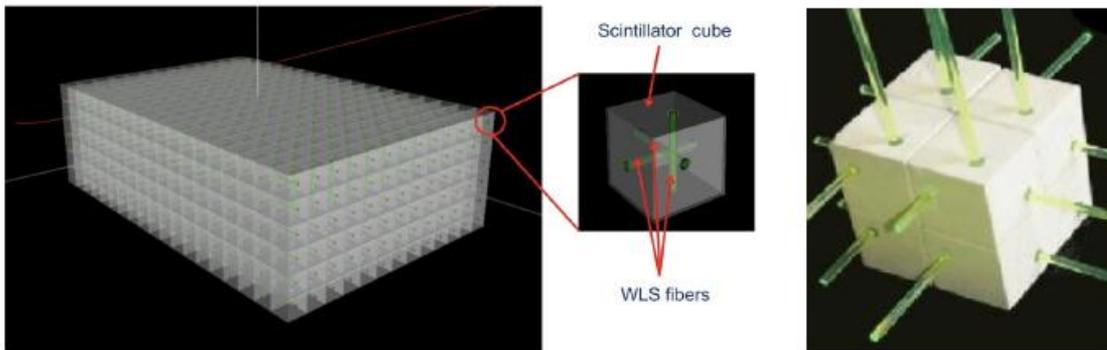
BSM = other BSM physics, historically more phenomenologically-oriented.  
Steriles, NSI, dark matter, BSM searches at the ND.



- Long baseline neutrino oscillation DUNE for CPV phase and mass ordering measurements.
- Staged year one is 2026 with neutrino beam, ND, and FD ready.
- Expected to produce a variety of new physics based on different types of interactions and different target materials.
- Korean contributions for protoDUNE, ND/3DST detector, Neutron study for reconstruction of anti muon neutrinos.
- Plan to participate protoDUNE analysis.
- Both global and local activities are waiting for participation of young researchers and students .
- 중앙대 중성미자 연구실  
중성미자 상호작용, Sim/Reco, 데이터 분야 연구원 채용 예정

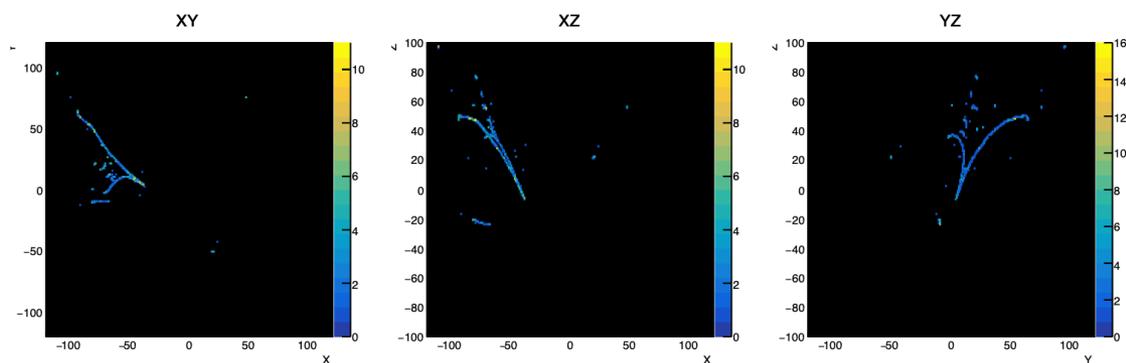
- Plastic scintillator detector with 1 cm x 1 cm x 1cm cubes      1.5 cm x 1.5 cm x 1.5 cm
- Light collected by 3 wavelength shifting fibers
- Each cube etched chemically to keep light entrapped inside the cube
- Read out by MPPC at 3 faces
- $4\pi$  coverage, 300 MeV/c proton threshold, 0.5 ns timing for MIP

Sunwoo Gwon  
for KPS 2020F

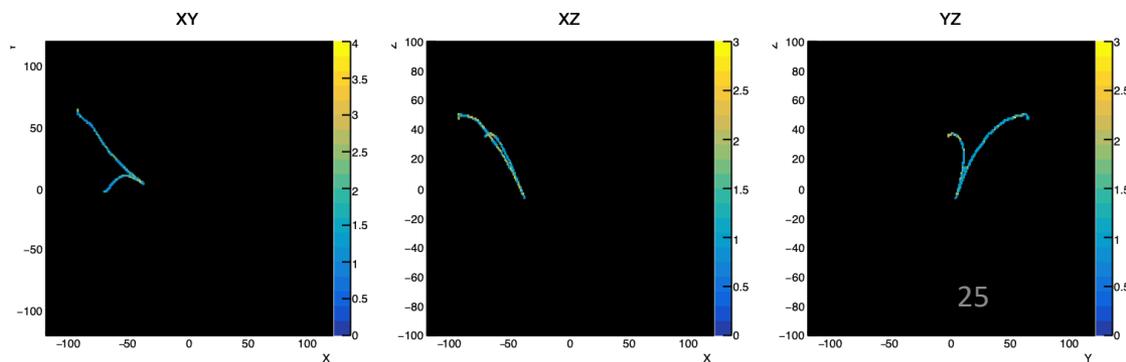


## Event Display

contains all neutron,  
gamma induced hits



only the cluster



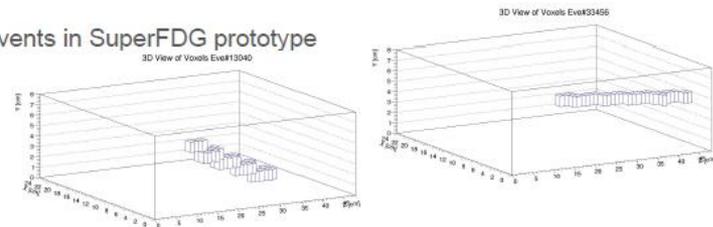
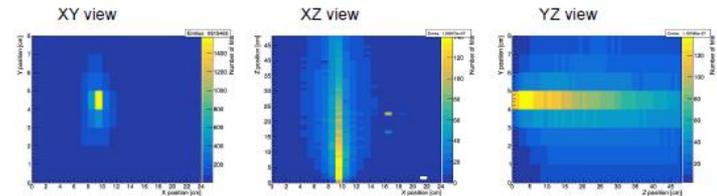
# Los Alamos Neutron Beam test

- SuperFGD 24x8x48 (2019, 2020)
- 3DST prototype 8x8x32 (2020)



## Neutron beam data

- Neutron beam observed in SuperFGD prototype detector
- Neutron events in SuperFGD prototype detector



## Joint T2K-DUNE 3D Scintillator R&D Group Institutions

CERN

Louisiana State University, USA

University of Pittsburgh, USA

Stony Brook University, USA

ETH Zurich, Switzerland

University of Pennsylvania, USA

High Energy Accelerator Research Organization (KEK), Japan

South Dakota School of Mines and Technology, USA



University of Geneva, Switzerland

Imperial college, UK

University of Rochester, USA

University of Tokyo, Japan

Chung-Ang University, South Korea

